AN EXAMINATION OF THE SUNSPOT AREAL DATASET, 1875–2017: PAPER II, HEMISPHERIC DIFFERENCES

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ABSTRACT

This is the second paper in an anticipated three-part study of the sunspot areal dataset. Examined are the annual variations of the northern (N) and southern (S) hemispheric sunspot area (SSA), number of active region entries (NARE), and the mean area per entry (MAE) for the interval 1875–2017, spanning solar cycles (SCs) SC12–SC24. For the overall interval of 1875– 2019, SSA(N) has been larger than SSA(S) for 77 of the 145 years. Likewise, for SC12–SC24, SSA(N) has been greater than SSA(S) during the ascending phase of the solar cycle (i.e., 35 of 52 years), whether the SC is an even- or odd-numbered SC, while SSA(S) has been greater than SSA(N) during the descending phase of the solar cycle (51 of 93 years). Minimum SSA(N) and SSA(S) have occurred in the same year in only 7 of 13 SCs, including SC14, SC16 and SC19– SC23. Maximum SSA(N) and SSA(S) have occurred in the same year only once (in SC15). Maximum SSA(S) usually occurs after maximum SSA(N), true for 9 of 13 SCs. Maximum SSA(S) preceded maximum SSA(N) in SC16, SC18, and SC19. Multiple peaks in SSA(N) or SSA(S), typically 2–3 years apart, have often been seen (e.g., SC12–SC16 and SC18–SC22). SC12 had the smallest maximum SSA(N) (452.0 millionths of a solar hemisphere), while SC19 had the largest maximum SSA(N) (2,277.2 millionths of a solar hemisphere). SC14 had the smallest SSA(S) (600.3 millionths of a solar hemisphere), while SC18 had the largest SSA(S) (1,642.9 millionths of a solar hemisphere). The average time from minimum to maximum SSA(N) is 3.8 years (range 2–5 years), while it is 4.8 years for SSA(S) (range 3–6 years). The average time from maximum $SSA(N)$ to the following minimum $SSA(N)$ is 6.9 years (range 5–9 years), while it is 6.2 years for SSA(S) (range 5–8 years). The largest N-S asymmetry coefficients for SSA occurs between –3 and +1 years about sunspot minimum. The largest N-S asymmetry coefficient in SSA occurred in 2019 (SC24). The largest area sunspot occurred in 1947 (SC18) in the southern hemisphere and measured 6,132 millionths of a solar hemisphere, nearly twice the size of the average maxima of the largest area spots in the other 12 SCs. SSA and NARE are highly correlated with each other and with SSN. Minimum MAE(N) and MAE(S) may have occurred, respectively, in 2018 and 2019, highly suggestive that SSN(m) for SC25 either occurred in 2019 or will occur in 2020.

INTRODUCTION

This is the second paper in an anticipated three-part study of the sunspot areal dataset, 1875–2017. Paper I (Wilson 2020) provided a general overview of the sunspot areal dataset, examining—in particular—the annual variations of selected parameters and inferred correlations that are apparent in the sunspot areal dataset. This paper (Paper II) examines the hemispheric

differences of selected parameters using the sunspot areal database. Paper III will examine annual variations of the magnetic complexity of sunspots.

 North-South (N-S) asymmetry has long-been recognized by solar observers as related to the distributions of major flare events, sunspot magnetic classes and sunspot areas. For example, Newton and Milson (1955) studied the N-S asymmetry of sunspot areas during the interval 1874–1954 and concluded that the fluctuations are real, their work following previous work reported by Newcomb (1901), Maunder (1922) and Kiepenheuer (1953). Later, Waldmeier (1971) investigated the N-S asymmetry of sunspots during the interval 1874–1969, arguing that the real asymmetry was strengthened by a phase difference of the two hemispheres, where the phase shift is subject to a long period that contains eight 11-year SCs. Roy (1977) noted that magnetically complex sunspot groups displayed a more pronounced asymmetry during the interval 1962–1970 than noncomplex groups. Antalová and Gnevyshev (1983) found that an N-S sunspot area asymmetry was a persistent feature, especially for SC14–SC20. Swinson et al. (1986) noted that, in general, northern hemispheric activity peaked about 2 years after sunspot minimum, with even-numbered cycles having a greater peak in northern hemispheric activity. Vizoso and Ballester (1987) examined the N-S asymmetry in sudden disappearances of solar prominences, noting that it does not appear to be in phase with the solar cycle, instead peaking about the time of solar minimum and reversing in sign during solar maximum. Garcia (1990) noted that the N-S distribution of large flares appears periodic and approximately in phase with the solar cycle, with the most intense large flares showing the largest N-S asymmetry. Vizoso and Ballester (1990) performed an exhaustive study of the N-S asymmetry of sunspots during the interval 1874–1976 and found: (1) that the N-S asymmetry is statistically significant, (2) the highest values of the N-S asymmetry coefficient for sunspots are obtained around sunspot minimum, and (3) there is a long-term periodic behavior of about eight cycles in which the activity in one hemisphere is more important during the ascending branch while during the descending branch the activity becomes more important in the opposite hemisphere. Through the years, many other studies have followed, investigating the N-S asymmetry of sunspots/solar flares including, in part: Schlamminger (1991); Yi (1992); Carbonell et al. (1993); Oliver and Ballester (1994); Duchlev and Dermendjiev (1996); Watari (1996); Ataç and Özgüç (1996); Vernova et al. (2002); Temmer et al. (2002, 2006); Li, Wang, Xiong, et al., (2002); Li, Chen, Zhan, Shi et al. (2009); Li, Gao, Zhan (2008); Li, Gao, Zhan et al. (2009); Knaack et al. (2004); Joshi and Joshi (2004); Vernova et al. (2004); Ballester et al. (2005); Zharkov et al. (2005); Zolotova and Ponyavin (2006); Zolotova and Ponyavin (2007); Chang (2007, 2008); Li (2009); Donner and Thiel (2007); Zolotova et al. (2009); Badalyan and Obridko (2011); and Chowdhury et al. (2013).

In this Paper (II), several issues regarding the asymmetry of sunspots are examined, including: (1) the variation of the N and S hemispheric annual SSAs for the interval 1875–2017 in relation to SSN minimum (m) and maximum (M) occurrences, (2) the variation of the N- and S-hemispheric annual NARE, (3) the variation of N- and S-hemispheric annual MAE, (4) the annual asymmetry coefficients for SSA and NARE, and (5) the results of epoch analyses for these parameters based on SSN(m) occurrence.

METHODS AND MATERIALS

Two primary data sources are used in this study: (1) annual values of SSN, available online at http://sidc.oma.be/silo/datafiles and (2) annual values of SSA, available online at http://solarcyclescience.com/activeregions.html. Other parameters taken from the SSA dataset (Greenwich Observatory (RGO) interval 1875–1976) and the United States Air Force/National Oceanic and Atmospheric Administration (USAF/NOAA) interval 1977–2019) include (1) NARE and (2) MAE, where MAE is computed as SSA divided by NARE times the number of days inthe year.

RESULTS AND DISCUSSION

Table 1 provides the basic data used in this investigation, spanning the interval 1875– 2017. The yearly 2018 and 2019 values are also shown but have not been included in the parametric means, which are given at the bottom of the table, both in terms of overall means (and standard deviations, *sd*) and even- and odd-numbered sunspot cycle means (and *sd*). Given in table 1 are the (1) NARE, (2) SSA, (3) MAE, and (4) asymmetry for NARE and SSA. For NARE, SSA, and MAE, the yearly N hemispheric, S hemispheric, and combined (C) hemispheric parametric values are given. Also given is each cycle's largest observed area (active region), denoted LAAR, and the hemisphere (H) in which it occurred. Also identified are the m and M parametric values for each SC, SC12–SC24.

@ means end of the RGO dataset.

The largest individual spot group was active region 1488603, having a corrected area spot size of 6,132 millionth of a solar hemisphere on 04/08/1947 and designated RGO type 8.

Table 2 provides a simple comparison of the timing occurrences of m and M for N and S parametric values with respect to the occurrences of SSN(m) and SSN(M), where the value '0' denotes that the parametric values occurred simultaneously (same year) with the corresponding SSN(m) and SSN(M) occurrences. Positive numbers indicate that the parametric values occurred (in years) after the corresponding SSN(m) and SSN(M) occurrences, while negative numbers indicate that the parametric values occurred (in years) before the corresponding SSN(m) and SSN(M) occurrences. Also given are the ascent (ASC) and descent (DES) intervals (in years) for each SC, where ASC is the time from to SSN(m) to SSN(M) for cycle *n*, while DES is the time from SSN(M) for cycle *n* to SSN(m) for cycle $n + 1$.

Note: # means SC24 DES is unknown; assumes SC25 in 2019.

Figures 1–3 display the annual variations of (1) (a) $SSA(N)$ and (b) $SSA(S)$, (2) (a) NARE(N) and (b) NARE(S), and (3) (a) MAE(N) and (b) MAE(S), respectively. Across the top of figures 1 and 2 are identified the years when SSN(m) and SSN(M) occurred per SC (located along the bottom in fig. 3), where SSN(m) occurrences are shown as unfilled triangles and SSN(M) occurrences are shown as filled triangles. The data intervals are divided according to whether the values come from the RGO record (1875–1976) or the extended record based on the USAF/NOAA observations (1977–present). The horizontal line in each figure gives the overall mean (given to the right along with *sd*) for the interval 1875–2017 and the numbers SC12–SC24 refer to the specific SCs. Plainly, the years 2018/2019 represent transitional years between the ending of SC24 and the onset of SC25, which is anticipated to occur either in late 2019 or 2020,

based on smoothed monthly mean SSN (cf. Wilson 2019a, 2019b, 2020). (It is now known that SC25 minimums for SSN and SSA occurred in 2019.)

Figure 1. Annual variation of (a) SSA(N) and (b)SSA(S).

Figure 2. Annual variation of (a) NARE(N) and (b) NARE(S).

Figure 3. Annual variation of (a) MAE(N) and (b) MAE(S).

Figure 1a shows that the maximum in SSA(N) per SC increased between SC12 and SC19 and decreased between SC19 and SC24. SC19 had the largest yearly SSA(N), measuring 2,277.2 millionths of a solar hemisphere and occurring in 1959, some 2 years after SC19 SSN(M). The timing of the maximum for SSA(N) has occurred simultaneously with the maximum SSN(M) for only 6 of 13 SCs. For SC12, SC13, SC20, and SC24, maximum SSA(N) preceded SSN(M), while for SC18, SC19 and SC23, maximum SSA(N) followed SSN(M). Minimum SSA(N) nearly always has occurred simultaneously with SSN(m), with only SC12, 17, and 21, having minimum SSA(N) occurrences that differed by 1 year from SSN(m).

Figure 1b shows that the maximum in SSA(S) per SC also increased, especially, between SC14 and SC18–SC19 and decreased afterwards through SC24. The largest yearly SSA(S) measured 1,642.9 millionths of a solar hemisphere and occurred in 1947 (SC18), simultaneously with SC18 SSN(M). The timing of maximum in SSA(S) is found to have followed SSN(M) in 7 of 13 SCs, occurring simultaneously in 5 of 13 SCs and preceded by one year only in SC16. Like SSA (N), the minimum in SSA(S) usually occurs simultaneously with SSN(m), with only SC14,

SC16, SC17, and SC21 having minimum SSA(S) occurrences that differed by one year from $SSN(m)$.

Comparing the minimums of SSA(N) and SSA(S), they occurred simultaneously in 7 of 13 SCs, including SC14, SC16 and SC19–SC23. Comparing the maximums of SSA(N) and SSA(S), they have occurred simultaneously in only 1 of 13 SCs (SC15). The maximum SSA(S) usually occurs after the maximum SSA(N), true for 9 of 13 SCs. For SC16, SC18, and SC19, the maximum SSA(S) occurred earlier than the maximum SSA(N). The mean SSA(N) measures 425.0 millionths of a solar hemisphere (*sd* = 399.5 millionths of a solar hemisphere), while the mean SSA(S) measures 412.5 millionths of a solar hemisphere ($sd = 381.4$ millionths of a solar hemisphere); hence, there is no statistically significant difference in the hemispheric means based on SSA.

Interesting is that many SCs have multiple peaks in SSA(N) and SSA(S), typically 2–3 years apart. Such behavior is clearly seen in SC12, SC13, SC16, SC18, SC19, and SC21 in SSA(N) and SC14, SC15, SC20, and SC22 in SSA(S). The rankings of maximum SSA(N) from least to greatest are as follows: SC12 (452.0); SC24 (561.4); SC13 (602.0); SC16 (725.9); SC14 (749.3); SC15 (853.6); SC23 (922.8); SC20 (1033.6); SC18 (1,180.2); SC21 (1,289.9); SC17 (1,316.1); SC22 (1,363.8); and SC19 (2,277.2). The rankingsof maximum SSA(S) from least to greatest are as follows: SC14 (600.3); SC15 (680.4); SC16 (681.9); SC20 (683.3); SC12 (820.9); SC24 (940.8); SC13 (944.0); SC23 (1,103.9); SC17 (1,129.2); SC21(1,227.1); SC22 (1,555.8); SC19 (1,608.8) and SC18 (1,642.9). On average, the maximum SSA(N) measures 1,025.2 millionths of a solar hemisphere ($sd = 482.1$) and the maximum SSA(S) measures 1,047.6 millionths of a solar hemisphere $\left(\frac{sd}{270.2}\right)$. The average time from minimum to maximum SSA(N) is 3.8 years (range 2–5 years), while the average time from minimum to maximum SSA(S) is 4.8 years (range 3–6 years). The average time from maximum to minimum SSA(N) is 6.9 years (range 5–9 years), while the average time from maximum to minimum SSA(S) is 6.2 years (range 5–8 years). For SC24, the elapsed time from its maximum to minimum SSA(N) for SC25 measures at least 6 years (if 2018 marks the occurrence of the minimum of SSA(N) for SC25), while the elapsed time from maximum to minimum SSA(S) for SC25 measures, at least 5 years (presuming 2019 marks the occurrence of minimum SSA(S) for SC25).

Figure 2 shows the yearly hemispheric variation of NARE, following the same format used in figure 1. Regarding figure 2a, it shows that the maximum in NARE(N) per SC increased between SC12 and SC19 and decreased between SC19 and SC24, as it occurred with SSA(N). SC19 had the largest yearly NARE(N), numbering 3,173 entries and occurring in 1959, some 2 years after SC19 SSN(M), as also was seen for SSA(N). The timing of the maximum for NARE(N) has occurred simultaneously with SSN(M) for only 6 of 13 SCs. For SC12, SC20, and SC24, the timing of maximum NARE(N) preceded SSN(M), while for SC18, SC19, and SC23, maximum NARE(N) followed SSN(M). Minimum NARE(N) nearly always has occurred simultaneously with $SSN(m)$, with only $SC12$, 15, and 17 having minimum NARE(N) occurrences that differed by one year from SSN(m). Regarding figure 2b, it shows that the maximum in NARE(S) per SC also increased, especially, between SC14 and SC18–SC22 and decreased afterwards through SC24. The largest yearly NARE(S) numbered 2,917 entries and occurred in 1991 (SC22). The timing of the maximum of NARE(S) is found to have followed SSN(M) in 6 of 13 SCs, occurred simultaneously in 5 of 13 SCs and preceded by one year only in SC16 and 20. Like NARE(N), the minimum in NARE(S) usually occurs simultaneously with SSN(m), with only SC12, SC15, and SC17 having minimum NARE(S) occurrences that differed by one year from SSN(m).

Comparing the minimums of NARE(N) and NARE(S), they occurred simultaneously in 5 of 13 SCs, including SC14, SC16, SC19, SC22, and SC23. Comparing maximums of NARE(N) and NARE(S), they have occurred simultaneously in only 4 of 13 SCs, including SC13, SC15, SC20, and SC21. The maximum NARE(S) usually occurs after the maximum NARE(N), true for 7 of 13 SCs, including SC12, SC14, SC17, SC19, and SC22–SC24. For SC16 and SC18, the maximum NARE(S) occurred earlier than the maximum NARE(N). The mean NARE(N) measures 873.3 entries ($sd = 703.6$), while the mean NARE(S) measures 852.7 entries ($sd =$ 680.6); hence, as with SSA, there is no statistically significant difference in the hemispheric means of NARE.

Double-peaking is found to have occurred in SC12, SC16, SC18, SC20 and SC21 for NARE(N) and in SC14, SC20, SC21 and SC23 for NARE(S), typically 2 years apart. The rankings of maximum NARE(N) from least to greatest are as follows: SC12 (895), SC24 (1,194), SC14 (1,235), SC13 (1,264), SC16 (1,301), SC23 (1,863), SC15 (1,938), SC17 (2,114), SC20 (2,159), SC18 (2,177), SC22 (2,532), SC21 (2,904) and SC19 (3,173). The rankings of maximum NARE(S) from least to greatest are is as follows: SC14 (1,053), SC12 (1,206), SC24 (1,251), SC16 (1,428), SC15 (1,572), SC20 (1,622), SC17 (1,711), SC13 (1,808), SC23 (1,940), SC18 (2,368), SC21 (2,534), SC19 (2,560) and SC22 (2,917). On average, the maximum NARE(N) numbers 1,903.8 entries ($sd = 703.2$), while on average, the maximum NARE(S) numbers 1,843.8 entries ($sd = 586.2$). The average time from minimum to maximum NARE(N) is 4.0 years (range 3–5 years), while the average for NARE(S) is 4.6 years (range 3–6 years). For NARE(N), six SCs had intervals of rise (minimum to maximum) of 3 years, including SC12, SC17, SC20–SC22, and SC24, while six SCs had intervals of rise of 5 years, including SC14– SC16, SC18, SC19, and SC23. For NARE(S), four SCs had intervals of rise of 4 years, including SC15, SC16, SC18, and SC21, while, four SCs had intervals of rise of 5 years, including SC17, SC20, SC22, and SC24 and three SCs had intervals of rise of six years, including SC12, SC14, and SC23. (Obviously, strong linear correlations exist between SSN, SSA, and NARE.)

Figure 3 shows the yearly hemispheric variation of MAE. For both hemispheres, MAE appears to have multiple peaks. Also, there appears to be a downward trend in MAE for both hemispheres between SC12 and SC15, followed by an upward trend between SC15 and SC18/19 and another downward trend between SC18/19 and SC24, one that is more pronounced in the N hemisphere than in the S hemisphere. The means for the two hemispheres are not statistically different.

A comparison of the minimum MAE(N) with SSN(m) (see table 2) reveals that 7 of 13 SCs had simultaneous occurrences (i.e., no lag), 3 had minimum MAE(N) prior (1–2 years) to $SSN(m)$, and 3 had minimum MAE(N) after (+1 year) $SSN(m)$. A comparison of minimum $MAE(S)$ with $SSN(m)$ reveals that 8 of 13 SCs had simultaneous occurrences, 4 had minimum MAE(S) prior $(1-2 \text{ years})$ to SSN(m), and 1 (SC24) had minimum MAE(S) after $(+1 \text{ year})$ SSN(m). Minimum MAE(N) and MAE(S) for SC25 may have occurred, respectively, in 2018 and 2019. Hence, SSN(m) for SC25 should be expected sometime about 2016–2019 based on the minimum occurrence of MAE(N) and about 2017–2020 based on the minimum occurrence of MAE(S). In terms of SSN, it decreased in value between 2016 and 2019, measuring 39.9 in 2016, 7.0 in 2018, and 3.6 in 2019. Clearly, SSN(m) for SC25 is near, expected to occur either in 2019 or 2020. (SSN(m) for SC25 occurred in 2019.)

The rankings from least to greatest maximum MAE(N) are as follows: SC24 (187), SC22 (197), SC13 (199), SC15 (208), SC12 (213), SC16 (218), SC20 (222), SC23 (222), SC14 (239), SC21 (249), SC19 (262), SC17 (267), and SC19 (293). In terms of rankings from least to greatest maximum MAE(S), it is as follows: SC20 (176), SC16 (199), SC22 (200), SC14 (208), SC15 (209), SC23 (212), SC24 (229), SC21 (231), SC13 (235), SC17 (241), SC19 (246), SC18 (253) and SC12 (265). The maximum MAE(N) tends to occur before the maximum MAE(S), true for 8 of 13 SCs. On average, the maximum MAE(N) measures 229 millionths of a solar hemisphere $sd = 31$ millionths of a solar hemisphere), and the maximum MAE(S) measures 223 millionths of a solar hemisphere (*sd* = 25 millionths of a solar hemisphere).

Figure 4 depicts the yearly variations of the asymmetry based on NARE (top) and SSA (bottom). Asymmetric values near zero simply mean that both hemispheres are approximately of equal parametric magnitude, whereas asymmetric values of large magnitude indicate that one hemisphere is of greater parametric magnitude than the other hemisphere. The greatest positiveor negative asymmetric values tend to occur near SSN(m). For Asymmetry (NARE), the maximum asymmetric value has occurred -3 to $+1$ years relative to SSN(m) occurrence. For Asymmetry (SSA), the maximum asymmetric value has occurred -3 to $+2$ years relative to SSN(m) occurrence. Large positive asymmetric values (the largest on record) occurred in 2019 for both NARE and SSA. Hence, one expects SSN(m) for SC25 to very probably occur very soon (i.e., sometime between 2019 and 2022, assuming the year 2019 marks the greatest asymmetric value year). For 11 of 13 SCs, maximum Asymmetry (SSA), ignoring sign, occurs either simultaneously with SSN(m) (SC12, SC13, SC17, and SC23) or before SSN(m)(SC14, SC15, SC18, SC19, SC21, SC22, and SC24).

Figure 4. Annual variation of (a) Asymmetry(NARE) and (b) NARE(SSA).

Table 3 gives the parametric m and M values and the mean and *sd* values for SC12–SC24 and for even- and odd-numbered cycles (SC24 values are incomplete). Figure 5 depicts the m parametric cyclic variations (top) for SC12–SC25 and the M parametric variations (bottom) for SC1–SC24 (i.e., using the combined hemispheres). Figure 6 shows the m (top) and M (bottom) parametric cyclic hemispheric variations for SC12–SC25 and SC12–SC24, respectively. The downward pointing arrows in figures 5 and 6 denote that the true m values for SC25 could be smaller than shown, especially if SSN(m) for SC25 occurs in 2020 or later. The largest parametric m values occurred in SC21 (fig. 5), while the largest parametric M values occurred (fig. 5) in SC18 (MAE), SC19 (SSN and SSA) and SC21 (NARE). The largest m value based on N hemispheric values occurred (fig. 6) in SC21 (SSA and MAE) and SC20 (NARE), while the largest m value based on southern hemispheric values occurred in SC23 (SSA), SC21 (NARE) and SC18 (MAE). The largest M values based on northern hemispheric values (Figure 6) occurred in SC19 (SSA and NARE) and SC18 (MAE), while the largest M values based on southern hemispheric values occurred in SC18 (SSA), SC22 (NARE) and SC12 (MAE).

Figure 5. Cyclic variation of (a) SSN(m), (b) SSN(M), (c) SSA(m), (d) SSA(M), (e) NARE (m), (f) NARE(M), (g) MAE(m), and (h) MAE(M).

Figure 6. N-S cyclic variation of (a) SSA(m), (b) SSA(M), (c) NARE(m), (d) NARE(M), (e) MAE(m), and (f) MAE(M).

Figure 7 displays the parametric variations based on SC12–SC23 relative to SSN(m) for (a) SSA, (b) NARE, and (c) MAE. In the figure, the thicker darker line represents the northern hemispheric parametric mean; the thinner lighter line represents the southern parametric mean; the darker dotted line (filled circles) represents SC24 northern hemispheric values; and the lighter dotted line (unfilled circles) represents SC24 southern hemispheric values. Marked across the top are the SSN(M) occurrences for SC12–SC24. The *t* = 10 values for SC24 correspond to the yearly 2018 values (see table 1; the yearly 2019 values are lower still for SSN, SSA, SSA(S), MAE(S), NARE, NARE(N) and NARE(S)). Interesting is the behavior of SSA(N), NARE(N), and MAE(N) near SSN(m) for SC24. Comparison of the 2018 yearly values to values in 2006– 2008 are suggestive that SC25(m) should be expected to occur within two years following 2018 (i.e., 2019 or 2020).

Figure 7. N-S variation of mean SC12–SC23 and annual SC24 values relative to SSN(m) for (a) SSA, (b) NARE and (c) MAE, for elapsed time *t* **(in years) equal to –5 to +10.**

Figure 8 is similar to figure 7 but now is plotted in terms of even- and odd-numbered SCs. For the odd-numbered SC, the plotted SC24 values presume SSN(m) for SC25 in 2019. If it turns out that the year 2020 marks the true occurrence of SSN(m) for SC25, then the plotted

Figure 8. N-S variation of mean even- and odd-numbered SC and annual SC24 values relative to SSN(m) for (a) SSA, (b) NARE and (c) MAE, for elapsed time *t* **(in years) equal to –5 to +10.**

Figure 9 plots Asymmetry(SSA) and Asymmetry(NARE). Again, SC24 values are plotted assuming SSN(m) for SC25 occurred in 2019. The greatest negative Asymmetry(SSA) for SC24 occurred at $t = -2$ (2006), while the greatest positive Asymmetry(SSA) for SC24 appears to have occurred at *t* = 11 (2019, not shown; see table 1), surpassing the value at $t = +3$ (2011).

Figure 9. Variation of Asymmetry of mean of even- and odd-numbered SC and annual SC24 values relative to SSN(m) for (a) SSA and (b) NARE, for elapsed time *t* **(in years) equal to –5 to +10.**

Figure 10 simply shows the cyclic variation of LAAR/H. The northern hemisphere had the largest spots at minimum (m) in SC12, SC14–SC17, and SC21 and the largest spots at maximum (M) in SC12, SC16, SC17 and SC19–SC22. For SC24 the S hemisphere had the largest LAAR/H at both m and M. The largest LAAR/H at maximum occurred in SC18 in 1947 (SC18), measuring 6,132 millionths of a solar hemisphere, while the LAAR/H at minimum occurred in 1934 (SC17), measuring 1,169 millionths of a solar hemisphere.

Figure 10. Cyclic variation of minimum and maximum LAAR/H.

Table 4 summarizes the behavioral dominance (i.e., number of years) of the N-S SSA per SC, both in terms of the overall SC and the ASC and DES intervals. For SC12–SC24, SSA(N) was larger than SSA(S), 75 versus 66 years. During the ASC interval, SSA(N) likewise, has been the more dominant, 35 versus 17 years, while during the DES interval, SSA(S) has been the more dominant, 49 versus 40 years. For even-numbered SCs, SSA(N) has been the more dominant overall and during the ASC interval, 43 versus 33 years, and 21 versus 9 years, respectively, while for odd-numbered SCs, SSA(S) has been the more dominant during the DES interval, 25 versus 18 years, while being less dominant during the ASC interval, 8 versus 14 years.

	Overall		ASC		DES	
Cycle	N	S	N	S	N	S
12	3	8	3	$\overline{2}$	0	6
13	3	9	$\overline{2}$	$\overline{2}$	$\overline{1}$	$\overline{7}$
14	5	$\overline{7}$	3	$\mathbf{1}$	$\overline{2}$	6
15	9	$\mathbf{1}$	4	0	5	$\mathbf{1}$
16	9	$\overline{1}$	4	$\mathbf{1}$	5	0
17	5	6	$\mathbf{1}$	3	4	3
18	5	5	$\mathbf{1}$	$\overline{2}$	4	3
19	$\overline{7}$	3	$\overline{2}$	$\mathbf{1}$	5	$\overline{2}$
20	8	4	4	0	4	4
21	4	6	3	0	$\mathbf{1}$	6
22	6	4	$\overline{2}$	$\overline{1}$	4	3
23	4	8	$\overline{2}$	$\overline{2}$	$\overline{2}$	6
24#	\geq 7	≥ 4	4	$\overline{2}$	\geq 3	\geq 2
Total#	≥ 75	≥ 66	35	17	≥ 40	≥ 49
Even#	≥ 43	≥ 33	21	9	\geq 22	\geq 24
odd	32	33	14	8	18	25

Table 4. N-S SSA dominance per SC.

Note: # means SC24 DES is unknown; excludes values for 2019.

In conclusion, this study has examined the N-S yearly variations of SSA, NARE, and MAE. For the overall interval of 1875–2019, the northern hemisphere has been the more dominant hemisphere in 77 of 145 years based on SSA. Likewise, for SC12–SC24, SSA(N) has been greater than SSA(S) during the ascending phase of the solar cycle (i.e., 35 of 52 years), whether the SC is an even- or odd-numbered SC, while SSA(S) has been greater than SSA(N) during the descending phase of the solar cycle (51 of 93 years). Minimum SSA(N) and SSA(S) have occurred in the same year in only 7 of 13 SCs including SC14, SC16, and SC19–SC23. Maximum SSA(N) and SSA(S) have occurred in the same year only once (in SC15). Maximum SSA(S) usually occurs after maximum SSA(N), true for 9 of 13 SCs. Maximum SSA(S) preceded maximum SSA(N) in SC16, SC18 and SC19. Multiple peaks in SSA(N) or SSA(S), typically 2–3 years apart, have often been seen (e.g., SC12–SC16 and SC18–SC22). SC12 had the smallest maximum SSA(N) (452.0 millionths of a solar hemisphere), while SC19 had the largest maximum SSA(N) (2,277.2 millionths of a solar hemisphere). SC14 had the smallest SSA(S) (600.3 millionths of a solar hemisphere), while SC18 had the largest SSA(S) (1,642.9 millionths of a solar hemisphere). The average time from minimum to maximum SSA(N) is 3.8 years (range 2–5 years), while it is 4.8 years for SSA(S) (range 3–6 years). The average time from maximum $SSA(N)$ to the following minimum $SSA(N)$ is 6.9 years (range 5–9 years), while it is 6.2 years for SSA(S) (range 5–8 years). The largest N-S asymmetry coefficients for SSA occurs between –3 and +1 year about sunspot minimum. The largest N-S asymmetry coefficient in SSA occurred in 2019 (SC24). The largest area sunspot occurred in 1947 (SC18), occurring in the southern hemisphere and measuring 6,132 millionths of a solar hemisphere, nearly twice the size of the average maxima of the largest area spots in the other 12 SCs. SSA and NARE are highly correlated with each other and with SSN. Minimum MAE(N) and MAE(S) may have occurred, respectively, in 2018 and 2019, highly suggestive that SSN(m) for SC25 either occurred in 2019 or will occur in 2020. (SSN(m) for SC25 is now known to have occurred in 2019.)

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